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# Optical Emission Spectroscopy of 2.45 Ghz Microwave Induced Plasma

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# Abstract

Optical emission spectroscopy measurements are performed to obtain the electron temperature and electron number density of argon plasma in 2.45 GHz microwave discharge for different filling pressures (0.4 - 0.7 mbar). Emission spectrum ranging from (380-710 nm) is recorded in axial and radial directions by using ocean spectrometer (HR4000). The electron temperature is determined from the slope of Boltzmann's plot that uses the intensity of several spectral lines versus their corresponding excitation energies. The electron number density is determined from the Stark broadening of well-isolated Ar-I (696.54 nm) spectral line after de-convolution of the Voigt profile. It is observed that electron temperature and electron number density is higher when optical emission is recorded in axial direction in comparison with radial direction. This fact may be ascribed to energy loss of electromagnetic wave as it propagates along the plasma column, which makes the plasma parameters to vary continuously along the discharge region.

**Keywords:** Microwave induced plasma, optical emission spectroscopy, plasma parameters, Stark-broadening.

### Introduction

Glow discharges generated by microwave power supplies through coupling of electromagnetic radiation of frequency ranging from 300 MHz to 10 GHz are named as microwave induced plasmas. Power supplies operating at a frequency of 2.45 GHz are readily available due to their common use in industry and home heating applications, which makes suitable power source for microwave plasma generation. This kind of plasma may be considered in local thermodynamic equilibrium (LTE) and non-local thermodynamic equilibrium (non-LTE) depending on electron number density and plasma pressure [Bogaerts et al., 2002, Zhan et al., 2003, Fantz et al., 2006]. For low electron density plasma only a few collisions occur and resultantly plasma species are characterized by different temperatures due to insufficient collisional transfer of energy among them. Thus electron number density plays an important role in differentiating the

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LTE and non-LTE plasma regimes. The potential applications of non-LTE plasmas are etching, cleaning and deposition of thin films [Zhan et al., 2003]. Among the plasma species, electrons have higher thermal velocities due to their lighter mass and therefore can be easily accelerated by the electromagnetic radiation as compared to other plasma species. These energetic electrons in turn may excite or ionize the plasma species having energy of several electron volts above their ground state by inelastic collisions. Therefore, electron temperature [Te] and electron number density [ne] are important parameters for describing the plasma discharges as a function of discharge conditions. Moreover, in many cases, the rare gases are used as a feed gas or added as a trace gas in low pressure discharges, and the intensity of their optical emission is used to determine these parameters intrusively [Fantz et al., 2006, Abrar et al., 2013, Ahmad et al., 2009]. Especially, argon mixed reactive plasmas are widely used in various material processing applications [Zhang et al., 2000, Qayyum et al., 2005, Behringer et al., 1991, Boffard et al., 2004]. In plasma processing, elementary processes such as ionization, dissociation, and excitation as well as ion bombardment are closely related to the plasma parameters and therefore, plasma diagnostics are important in understanding the mechanism material processing applications of [Malyshev et al., 1997, Zhu et al., 2010, Czech et al., 2011, Bibinov et al., 2008, Zhu et al., 2008].

Argon plasma has been widely used for many kinds of material processing and surface modification applications, because argon is a kind of inert gas and has an appropriate atomic mass amongst noble gases. Knowledge of plasma parameters, such as electron temperature, electron number density and EEDF play a key role in the understanding of physics and chemistry of argon plasma for better optimization of processing reactor [Ahmad et al., 2009]. A lot of research work on microwave induced plasmas has been carried out so far to establish correlation between the plasma and discharge parameters conditions

[Konjevi et al., 2009, Mattei et al., 2011, Zhu et al., 2008]. However, a little investigation is carried out on correlation of plasma parameters with discharge conditions in microwave induced argon plasma via spectroscopy diagnostics.

Optical emission spectroscopy is an appropriate and inventive technique to determine the plasma parameters such as electron temperature, electron number density and EEDF [Qayyum et al., 2005, Czech et al., 2011]. The electron temperature and the electron number density in low pressure plasma may be extracted from the emission spectrum of atomic and molecular species that are excited by direct electron impact process and depopulated by the spontaneous emission of radiation [Bibinov et al., 2008, Zhu et al., 2008]. Several spectroscopic methods are being used to determine the electron temperature [T<sub>e</sub>] including line-ratio method and Boltzmann's plot method. However, Boltzmann's method is preferred because it uses several spectral lines instead of two spectral lines and thus improves the accuracy of the measurements. This method adopted here to determine the electron density [n<sub>e</sub>] from the recorded single symmetrical profile is the Voigt fit on the shape and may be useful where no other diagnostic methods are available. The fit to Voigt profile is more appropriate, because it includes Doppler broadening as well as the Lorentzian profile. The Lorentzian profile has its roots in the Stark effect and is a consequence of radiating atoms interaction with free electrons and ions in the neighboring plasma region through local electric fields [Torres et al., 2007]. This profile can be used if the shape of the measured line is symmetric, which is usually the case of ionized emitters, where it is normally possible to neglect the interactions between the emitter and perturbing ions. In addition, in case of astrophysical plasmas, where no other diagnostic method is available, this method can be extremely helpful [Milosavljevic et al., 2001].

This paper reports the axial and radial measurements of the electron temperature and electron number density of argon plasma

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

in 2.45 GHz discharge versus filling pressure at fixed input power and gas flow rate. Boltzmann's plot method is used to determine the electron temperature of the argon plasma. The electron number density is calculated from the Stark broadening of Ar-I (696.54 nm) spectral line through Voigt profile fit after subtracting contribution of mechanisms. other broadening The measured electron temperature and electron number density are also discussed with reference to different electron and ionization processes involved in optical emission at different filling pressures.

## Experiment

The schematic of the experimental arrangement is shown in figure 1. It consists of a cylindrical stainless steel vacuum chamber having diameter of 26 cm and height of 60 cm with several multi-purpose ports. One port of the vacuum chamber is coupled to pumping system that comprises of rotary pump and diffusion pump. After proper evacuation argon is fed as a working gas and plasma is generated at different filling pressures (0.4 - 0.7 mbar) for spectroscopic analysis The microwave discharge is maintained at the frequency of 2.45 GHz by means of 800 watt microwave generator. The main components of the microwave generator are the high voltage system and magnetron tube. The high voltage system consists of step-up transformer and rectifier circuit for increasing input voltage from 220V to 2kV and then converting it from AC to DC. The magnetron operation is mainly based on the motion of electrons

under the combined effect of electric and magnetic fields. Therefore, the trajectory of the electrons is governed by two fundamental facts. Firstly, the force exerted by an electric field on an electron is proportional to the strength of the field. Secondly, the force exerted on an electron in a magnetic field is at right angle to the both field itself and the path of the electrons. Consequently, the trajectory of the electron under the Lorentz force from the cathode to the anode is curving rather than straight path. Permanent magnets are added above and below the magnetron tube to produce perpendicular magnetic field and resultantly electrons motion is circular. The electrons are emitted from the heated filament which also serves as the cathode of the magnetron tube. These electrons are accelerated from the cathode towards anode. In this way work is done on the electrons and therefore, energy from the power supply is given to them. As these electrons move toward a point where excess of negative charge already exist due to the presence of resonant cavities. So they are pushed back around the cavity, imparting energy to the oscillations at the natural frequency of the cavity. These driven oscillations of the charges around the cavities leads to radiation of electromagnetic waves: the output of the magnetron. Thus DC power is converted to the RF energy due to resonant cavities that serve as a tuned circuit. The output RF energy is transmitted by an antenna that is connected to the anode and extends into one of the tuned cavities. The antenna is coupled to the waveguide, a hollow metal enclosure; which transmit the RF energy to the plasma column.

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302



In this experiment, the microwave radiation are coupled with the chamber through a circular waveguide having diameter of 12.4 cm and length of 11 cm. Argon is injected at the top of the chamber in a controlled manner. Filling pressure in the chamber is monitored by using Pirani gauge. A quartz window is used to collect and thus analyze optical emission from the plasma under the different discharge conditions through the optical fiber with ocean spectrometer HR4000 having grating of 300 lines/nm and optical resolution of 0.75 nm (FWHM). The spectrometer can be adjusted both, in radial and axial direction to the plasma chamber for characterizing and co-relating the plasma parameters. The optical emission from the plasma is collected by placing the optical fiber at the centre of the top window as well as at the centre of side window. The exact location of the optical fiber used for collecting emission data is presented in the schematic of experiment. A typical spectrum of argon plasma at pressure of 0.6 mbar is

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

shown in figure. 2. The spectral lines in the emission spectra are identified and labeled

by using NIST data.



Figure 2: Typical emission spectrum at pressure of 0.6 mbar and microwave power of 800 W.Spectroscopic Measurements

#### **Measurement of Electron Temperature**

In microwave plasma, the electrons gain energy from the electromagnetic radiation and transfer a part of their energy to the gas particles through collisions. These collisions may be elastic or inelastic depending upon the energy of incident electrons. Inelastic collisions promote some of the atoms to higher energy states that subsequently decay and emit characteristic photons. The intensity of these characteristic photons provides information of upper state population density of the emitting species and also electron energy distribution subjected to excitation by electron impact process and de-excitation by spontaneous emission [Bibinov et al., 2008, Zhu et al., 2008]. Therefore, the electron temperature can be obtained from the slope of the Boltzmann's plot by assuming that EEDF is Maxwellian and atoms are excited from their ground state without any multi-stage excitation or de-excitation. Explicitly the relation is [Abrar et al., 2013, Qayyum et al., 2007, Naveed et al., 2006].

$$ln\left(\frac{I_{ki}\lambda_{ki}}{A_{ki}g_k}\right) = \frac{-E_k}{k_BT} + C \tag{1}$$

Where  $\lambda_{ki}$  is the wavelength in (nm),  $I_{ki}$  is the measured intensity in (a.u),  $A_{ki}$  is the transition probability~  $10^8 \text{sec}^{-1}$ ,  $g_k$  is the statistical weight and C is the constant for a given atomic species. Boltzmann's plot method reduces the errors arising from the non-LTE plasma since several spectral lines of different excitation energies are used in this method. The electron temperature is measured as a function of filling pressure by measuring the line intensity of number of argon lines from the spectrum. The measured intensities are then normalized for the spectral response of the spectrometer provided by the manufacturer. A graph is plotted for different values of  $ln\left(\frac{I_{kl}\lambda_{kl}}{A_{kl}g_k}\right)$ verses the energies of the upper level for each considered transition as shown in figure 3. The slope of the plot which is sensitive to the emission intensity of the spectral lines together with transition probability gives the electron temperature. The spectroscopic data of the spectral lines used in the Boltzmann's plot are presented in the table.

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

| λ(nm)  | Transition  | A <sub>ki</sub> (10 <sup>6</sup> s <sup>-1</sup> ) | g <sub>k</sub> | E <sub>k</sub> (cm <sup>-1</sup> ) |
|--------|---|--|----------------|------------------------------------|
| 340.61 | 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>1/2</sub> )4s - 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>1/2</sub> )7p | 0.39   | 1              | 124749.80                          |
| 367.06 | 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>1/2</sub> )4s - 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>1/2</sub> )6p | 0.31   | 5              | 122635.05                          |
| 405.45 | 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>3/2</sub> )4s - 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>1/2</sub> )5p | 0.27   | 3              | 118407.43                          |
| 425.11 | 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>3/2</sub> )4s - 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>3/2</sub> )5p | 0.11   | 3              | 116659.99                          |
| 549.01 | 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>3/2</sub> )4p - 3s <sup>2</sup> 3p <sup>5</sup> ( <sup>2</sup> P <sup>0</sup> <sub>3/2</sub> )6d | 0.85   | 5              | 123826.75                          |



Figure 3: Boltzmann plot for Ar-I spectral lines having different threshold excitation energies at filling pressure of 0.6 mbar.



Fig. 4: Electron temperature as a function of filling pressures for (a) axial (b) radial measurements

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

#### **Measurement of Electron Density**

The final shape and width of a spectral line is affected by different mechanisms including natural broadening, Doppler broadening, Stark broadening and instrumental broadening [Brugeat et al., 2004, Oavvum et al., 2006]. However, the broadenings like van der Waals broadening (due to dipolar forces high pressure discharges), natural in broadening (as a result of the uncertainty principle) and resonant broadening (as a consequence of electric dipole transition of the same species of atoms) are less important in this case and can be neglected

[Torres et al., 2007]. But, the Stark broadening that is proportional to the local electric fields produced by charge particles, particularly electrons surrounding the emitting atoms, as a function of their density and energy is important and thus eases the measurement of the electron density and field strength. Therefore, Stark broadening is one of the most important features of spectral line broadening that provides the electron density with the known electron temperature of plasma [Qayyum et al., 2006]. An experimentally measured emission profile of Ar-I (696.54 nm) for different filling pressures is presented in figure 5.



Figure 5: Emission profile of Ar-I (696.54 nm) spectral line for axial measurements at different filling pressures

The procedure of separating Stark broadened contribution from the total emission profile is detailed here. Generally the instrumental broadening, natural broadening and the Doppler broadening results into the Gaussian profile. Whereas by considering the electron broadening and ignoring the impact dynamics of ions during the radiative process, Stark broadening can result into Lorentzian profile. Thus the convolution of the Gaussian profile and the Lorentzian profile results in the so-called Voigt profile. Therefore, in order to analyze the experimental profile, the fit to Voigt profile is more appropriate, because it includes both

Doppler broadening as well as Lorentzian profile .Stark broadened part can be obtained after de-convolution of total broadened profile (Voigt profile) as both being statistical independent processes. The intensity distribution of the spectral line may be expressed by Voigt function [Milosavljevic et al., 2001].

$$V(\lambda) = V_o + V_{max} W_L^2 \times \int_{-\infty}^{\infty} \frac{exp(-t^2)}{W_L^2 + 4 \left[ \lambda - \left( \lambda_o + \left( W_G / 2\sqrt{\ln 2} \right) t \right) \right]^2} dt$$

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

where  $V_o$  is the baseline (offset) and  $V_{max}$  is the maximum of intensity(intensity for  $\lambda = \lambda_o$ ),  $W_L$  Lorentzian width (FWHM) and  $W_G$ Gaussian Width (FWHM). Since Doppler and instrumental profiles are both Gaussian hence Doppler width  $\Delta \lambda_D$  and instrumental width  $\Delta \lambda_1$  are related with each other according to the following expression [Moussounda et al., 1987].

$$\Delta \lambda_G = \sqrt{\left( \Delta \lambda_D \right)^2 + \left( \Delta \lambda_I \right)^2}$$

The statistical velocity distribution of the atoms or molecules results into the Doppler broadening of the spectral line in which each radiating atom or molecule undergoes a shift in their apparent frequency or wavelength. The radiating plasma species have several different velocities and thus result in many shifts in corresponding wavelength; effect of each is to broaden the line. In the case of Maxwell distribution of velocities the Doppler broadening has a Gaussian distribution. The Doppler broadening of the spectral line at its FWHM is written as [Moussounda et al., 1987]

$$\Delta \lambda_{Doppler} = 7.156 \times 10^{-7} \lambda_0 \times \sqrt{\frac{T}{M}}$$

Here  $\lambda_0$  is the wavelength of spectral line (696.54 nm), T is the temperature in K and M is the mass of the emitting particle in atomic mass unit. Therefore, Stark broadening of Ar-I (696.54 nm) line is obtained through the de-convolution of Voigt profile by using non-linear curve fitting and is used to determine the electron density by using the following relation [Ohno et al., 2006, Griem, Plasma spectroscopy, 1964]:

$${}^{(3)}_{\Delta\lambda_{1/2}} = 2[I + I.75\alpha(I - 0.75\,r)]\,\omega \tag{5}$$

Where  $\omega$  is the electron impact half widths in nm,  $\alpha$  is the ion broadening parameter in nm, r is the mean distance between ions divided by the Debye length. The values of  $\omega$  and  $\alpha$  of the well resolved and isolated Ar-I line (696.54 nm) have been given in Griem [Plasma spectroscopy, 1964]. By substituting the values of  $\alpha$  and  $\omega$  in Equation (5), the following expression is derived.

$$ln n_e = 1.20 ln [\Delta \lambda_{1/2}] + 44.2476 - 0.60 ln [T_e]$$

Here  $T_e$  is the electron temperature in *K* and  $n_e$  the electron number density in cm<sup>-</sup>



Figure 6: Electron density as a function of filling pressures determined from Stark broadening of Ar-I (696.54 nm) line for (a) axial measurements (b) radial measurements

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

#### **Results and Discussion**

Figure 2 shows the emission spectrum of 2.45 GHz argon plasma at a filling pressure of 0.6 mbar and input power of 800 watt. The prominent spectral lines in the spectrum are identified and labeled using NIST data. The emission lines in the spectrum have different intensity distribution that may be attributed to different population of their respective emitting states. This rises from the fact that emitting states have different threshold excitations and thus involve different group of electrons for their impact excitation.

Fig. 3 shows the Boltzmann's plot that uses the emission intensity of several spectral lines of argon having different threshold excitation energies. The electron temperature is determined from the slope of the Boltzmann's plot at a filling pressure of 0.6 mbar and it comes to be 0.64 eV. The variation of electron temperature versus filling pressure is presented in Fig. 4 for both measurements, i.e. axial as well as radial directions. It can be seen from the graph that the value of the electron temperature for axial direction is higher than radial direction for all filling pressures. This fact can be explained as follows: The axial region is directly exposed to microwave electric field by waveguide through quartz window and thus field is stronger in this region. Consequently, energy absorbed by electrons and also the collision rate of electrons with other plasma species is higher due to direct heating of electrons by microwave electric field and thus higher electron-impact ionization. The higher electron temperature and consequently higher electron number density is expected as a result of electron impact ionization. The microwave electric field intensity decreases with increasing distance from the launcher towards the end of plasma column due to attenuation of microwave power. The energy of electrons received from the electric field decreases and thus less electron impact excitation and ionization in the region away from launcher. The other reason may be the recombination loss of electrons in the outer portion of the plasma that is cooler compared to inner

portion. The re-absorption is also higher in radial direction due to large number of neutrals available in the path of photons and thus attenuation of emission intensity. Additionally, the spectroscopic technique provides the line of sight average intensity of optical emission by the radiating atoms. Therefore the difference in plasma length and plasma quality owing to different plasma conditions and radiating atoms involved in optical emission in axial and radial measurements may be the other reason. The results also show that the electron temperature decreases with increasing filling pressures. This decrease in electron temperature with filling pressure may be attributed to increase in number of collisions and consequently efficient collisional transfer of energy from the electron to the other plasma species. Fig. 5 shows the dependencies of Ar-I (696.54 nm) spectral profile as a function of filling pressure for axial measurements. It can be seen from the figure that the broadening of spectral line is strongly affected by the filling pressure and adopting de-convolution thus after procedure described earlier; FWHM of Stark broadening is obtained and related to electron number density.

Fig. 6 presents the comparison of electron number density versus filling pressure for both axial and radial direction. The result shows that the electron number density is higher in axial direction as compared to radial direction. As mentioned earlier, this suggests that the electron impact ionization is higher due to strong microwave electric field in the launching region. It is also seen that the electron density decreases as we move farther away from the axis of plasma column. This electron density gradient in radial direction is due to the electron energy gradient [Rousseau et al., 2006].

Moreover, surface wave discharges (SWDs) are a particular class of high-frequency (HF) discharges, where the plasma column being sustained by the electromagnetic surface wave and plasma plays a dual role, acting simultaneously as a component that absorbs the microwave power carried by the wave

F. Deeba, A. Qayyum and Nasir Mahmood (2015), Journal of Research in Spectroscopy, DOI: 10.5171/2015.172302

and as a medium for wave propagation. So the surface-wave creates the plasma that it needs to propagate by dissipating the power it carries. Thus, the wave loses its energy as it propagates along the plasma column, which makes the plasma parameters to vary continuously along the discharge region. For instance, the electron density decreases from the wave launcher toward the end of the plasma column. In a collisional discharge, the absorption of microwave power is thus a function of the collisional frequency of the electrons with other plasma species and is therefore, dependent on the pressure in the discharge. The decrease in the electron density with filling pressure may be due to cooling of energetic electrons through collisional transfer of energy to other plasma species, and thus reduction in electron impact ionization.

# Conclusion

The electron temperature and the electron number density in 2.45 GHz microwave argon plasma are determined by using the optical emission spectroscopy technique at constant input power and flow rate. Basic motivation of this experiment is to explore the effect of filling pressure on wave propagation together with excitation and ionization process involved in optical emission. Experimental results indicate that the plasma parameters achieve higher values near the axis compared to end region of the plasma column. This fact may be attributed to energy loss of electromagnetic wave as it propagates along the plasma column, which makes the plasma parameters to vary continuously along the discharge region. The decrease in the energetic number of electrons owing to collisional transfer of energy from electrons to other plasma species at higher pressures and consequently the reduction of electron impact ionization may be the cause of the decrease in electron number density.

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